1	Stratospheric Ozone in the Last Glacial Maximum
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ABSTRACT

Using the Whole Atmosphere Community Climate Model version 6, stratospheric ozone
in the Last Glacial Maximum (LGM) is investigated. It is shown that, compared with preindustrial
(PI) times, LGM modeled stratospheric temperatures increase by up to 8 K, leading to faster ozone
destruction rates for gas phase reactions, especially via the Chapman mechanism. On the other
hand, stratospheric hydroxyl radical (OH) and nitrogen oxides (NO <sub>x</sub> ) concentrations decrease by
10-20%, which decreases catalytic ozone destruction, thereby decreasing ozone loss rates. The net
effect of these two compensating mechanisms in the upper stratosphere (above 15 hPa) is a
vertically-integrated 1-3 Dobson Unit (DU) decrease during the LGM. In the lower stratosphere
(tropopause to 15 hPa), changes in the stratospheric overturning circulation and resulting transport
dominate changes in ozone. Consistent with a weakening of the residual circulation in the LGM,
lower stratospheric ozone increases by 2-5 DU in the tropics and decreases by 5-10 DU in the
extratropics, but the latter is partly compensated by ozone increases due to a lower tropopause. It
is found that tropospheric ozone decreases by about 5 DU in the LGM versus PI. Combined
changes in stratospheric and tropospheric ozone lead to a decrease in total ozone column
everywhere except over the northeast of North America, equatorial Indian and west Pacific Oceans
and East Antarctica. Surface ultraviolet radiation in the LGM versus PI increases over the
Northern Hemisphere mid- and high-latitudes, especially over the ice caps, and over the Southern
Hemisphere near 60° S.

Key words: Stratospheric ozone, Last Glacial Maximum, ultraviolet radiation, WACCM6

#### 1. Introduction

As a key component of the Earth system, stratospheric ozone protects life on Earth from hazardous ultraviolet (UV) radiation, and largely drives tropospheric chemistry. Rohrer and Berresheim [2006] showed that tropospheric hydroxyl radical (OH) is linearly correlated with UV radiation based on 5 years of measurements in Germany. Murray et al. [2014] found that stratospheric ozone, via its impact on surface UV radiation, is an important factor controlling variability in tropospheric OH over glacial-interglacial periods in a chemistry-climate model. In addition, stratospheric ozone could impact tropospheric climate [e.g. Nowack et al. 2015; Chiodo and Polvani, 2017; Noda et al., 2017; Szopa et al., 2019]. Szopa et al. [2019] found that modeled global surface temperature in the Eocene differed by as much as 14%, depending on whether the stratospheric ozone column abundance was consistently calculated or prescribed at preindustrial levels.

During the LGM, the stratosphere, including stratospheric ozone, is expected to be different from the preindustrial climate because of lower greenhouse gas concentrations (e.g., CO<sub>2</sub>, CH<sub>4</sub>, N<sub>2</sub>O), wide-spread ice sheets (up to 3-4km thick) in the northern hemisphere (NH) and lower sea surface temperatures (SSTs). Compared to the rich literature on stratospheric ozone for the current and future climates, there are only a few studies on stratospheric ozone in the Last Glacial Maximum (LGM) [Crutzen and Brühl, 1993; Martinerie et al., 1995; Rind et al., 2009]. From a lack of better knowledge, past model simulations of the glacial climate have often assumed that stratospheric ozone is similar to the preindustrial climate [e.g., Valdes et al., 2005; Kaplan et al., 2006; Levine et al., 2011]. However, as indicated above, there are multiple reasons to expect ozone in the LGM to be different to the preindustrial distribution, and this study provides an analysis of these differences using a state-of-the-art chemistry climate model.

Pioneering research about stratospheric ozone in the LGM was carried out by Crutzen and Brühl [1993], who used a "one and a half" dimensional (1½-D) model. Martinerie et al. [1995] utilized a two-dimensional (2-D) model to further investigate the chemical composition of the atmosphere during the LGM. Both of these studies found only small changes in stratospheric ozone between the LGM and the preindustrial climate, despite large changes in atmospheric greenhouse concentrations. Rind et al. [2009] examined stratospheric ozone in the LGM using the 3-D GISS Global Climate Middle Atmosphere model (4°x5° and 53 layers) with linearized ozone chemistry [McLinden et al. 2000]. The model used in Rind et al. [2009] did not consider stratospheric heterogeneous chemical processes nor tropospheric chemistry, but instead simulated tropospheric ozone using a simple relaxation approach. Rind et al. [2009] found that reduced nitrogen oxides (NO<sub>x</sub>) and chlorine in the LGM lead to increases in stratospheric ozone, and much higher ozone concentrations in the polar, lower stratosphere associated with an intensified Brewer-Dobson Circulation (BDC) in the lowermost stratosphere, which in turn leads to stronger stratosphere-to-troposphere transport of ozone.

Recent ice core observations of a proxy for tropospheric oxidants suggested that tropospheric ozone concentrations may have been higher in cold climates compared to warm climates [Geng et al., 2017], the opposite of expectations based on changes in temperature-dependent surface emissions of ozone precursor gases. Geng et al. [2017] hypothesized that higher tropospheric ozone in the extratropics could be due to increased transport of stratospheric ozone to the surface driven by an enhanced BDC in the glacial climate [Rind et al., 2001; Rind et al., 2009]. An alternative hypothesis was that reactive halogen chemistry could be enhanced in the glacial climate, with potential implications for tropospheric ozone abundances [Geng et al., 2017]. Using the Whole Atmosphere Community Climate Model version 6 (WACCM6), Fu et al. [2020]

showed that the BDC during the LGM is consistently slower than modern climate. By transporting stratospheric ozone from the tropics, where it is produced, to the poles, the BDC plays a substantial role in determining the spatial distribution of ozone and thus surface UV radiation. This study analyzes the changes in ozone and UV radiation in the LGM including those resulting from changes in BDC, as simulated by WACCM6 [Fu et al., 2020].

This paper consists of four sections. Section 2 provides a brief description of WACCM6 and the model simulations. Section 3 presents the main results, and section 4 summarizes and concludes the paper.

### 2. Model Description and Simulations

### 2.1 Model description

The WACCM6 is the high-top atmospheric component of the NCAR Community Earth System Model 2 [Gettelman et al. 2019]; our simulations have a horizontal resolution of 0.9° latitude by 1.25° longitude with 70 vertical levels extending from the surface to 6x10-6 hPa (~140 km). In WACCM6, stratospheric heterogeneous chemical processes are parameterized following Wegner et al. [2013] and Solomon et al. [2015]. Chemical kinetics and photochemical rate constants have been updated following JPL 2015 recommendations [Burkholder et al., 2015]. The updated tropospheric chemistry scheme in WACCM6 leads to an improvement to the isoprene oxidation as well as other ozone precursors, and thus the tropospheric ozone simulation, as compared to the observations [Emmons et al., 2020]. WACCM6 was run by specifying SST and sea ice. We coupled the WACCM6 model with the CLM4.0 land model [Oleson et al., 2010], and utilized the CLM4.0 LGM lower boundary conditions from the CESM Paleo Working group [e.g., Brady et al., 2013]. In the LGM simulations, the plant functional type distribution is fixed under

preindustrial conditions but the leaf area index, vegetation height, and biogenic emissions are estimated as a function of the changed climate [Thornton et al., 2007; Guenther et al., 2012].

WACCM6 is able to reproduce variability in the middle atmosphere as derived from reanalyses (e.g., MERRA, ERA-Interim), including the frequency of Stratospheric Sudden Warmings [Gettelman et al., 2019] and BDC [Fu et al., 2020]. WACCM6 has an internally generated Quasi-Biennial Oscillation in the lower stratosphere [Garcia and Richter, 2019]. The WACCM model suite has a realistic evolution of the Southern Hemisphere (SH) springtime ozone hole over the latter half of the 20th century [Solomon et al., 2015]. For more details, the readers are referred to Gettelman et al. [2019].

#### 2.2 Model simulations

Table 1 describes the simulations carried out explicitly for the study: 1) LGM simulation with prescribed SSTs derived from models (LGM<sub>PMIP3</sub>), 2) LGM simulation with SSTs based on proxy data (LGM<sub>PROXY</sub>), 3) preindustrial climate simulation (PI). With the first 10 years discarded for spin-up, the results presented here are based on the simulations from years 11 to 20; sensitivity studies extending simulations to 60 years find that using years 11-20 provides robust results for the analysis presented here.

The concentrations of CO<sub>2</sub>, CH<sub>4</sub>, and N<sub>2</sub>O for the LGM (PI) are 185 (284) ppm, 349 (810) ppb and 200 (273) ppb, respectively. The surface concentrations of long-lived atmospheric species represent the year 1850 for the PI, and all anthropogenic emissions are turned off in the LGM. The orbital parameters are conditions 21,000 years ago for LGM simulations and year 1850 for the PI simulation. For the PI simulation, the SSTs/sea ice are the climatological averages of the year 1870-1890 based on observations [Rayner et al., 2003] and include a seasonal cycle. For LGM<sub>PMIP3</sub>,

we used the PMIP3 multi-model mean SST/sea ice for 7 models that had both LGM and historical runs [Braconnot et al., 2012]. The SSTs and sea ice for the LGM were then obtained as the model differences between the LGM and modern climate simulations plus corresponding observed SSTs/sea ice in modern climate [Fu et al., 2020]. For LGM<sub>PROXY</sub>, the LGM-PI SST differences are taken from the MARGO proxy data [Kucera et al., 2005], while sea ice is the same as in LGM<sub>PMIP3</sub>. The LGM ice sheet topography was from Abe-Ouchi et al. [2015].

For ozone-depleting substances, WACCM6 considers 17 organic halogens, among which four (i.e., methyl bromide, methyl chloride, dibromomethane and bromoform) include natural sources. Therefore, we set methyl bromide, methyl chloride, dibromomethane and bromoform surface mixing ratios in the LGM to preindustrial levels, and while the remaining 13 species are not included in the LGM and PI simulations. Emissions of CO and NO<sub>x</sub> in the LGM were scaled with preindustrial emissions (Table 1) following Murray et al. [2014]. The CO emissions from fires, ocean and fuel combustion in the LGM were scaled to 0.1, 1 and 0 of the preindustrial levels while the NO<sub>x</sub> emissions from fires, soil and fuel combustion in the LGM were scaled to 0.1, 0.98 and 0 of the preindustrial values. The emission of NO<sub>x</sub> from lightning is simulated with the parameterization of Price and Rind [1992].

#### 3. Results

#### 3.1 Temperature changes

Figure 1 shows the annual- and zonal-mean temperature change for LGM<sub>PMIP3</sub> versus PI. Herein the difference between LGM and PI is referred to as the ice-age effect. Due to the ice-age effect, the upper stratosphere in the LGM warms by up to 8 K, mainly because of much lower greenhouse gas concentrations and hence reduced radiative cooling to space. Such temperature

changes are consistent with previous studies [Rind et al., 2001; Rind et al., 2009], and have a significant impact on stratospheric ozone concentration in the LGM, as will be shown later. The temperature changes in the troposphere are largely a response to the changes of SST and greenhouse gases, showing a maximum cooling of more than 4 K in the tropical upper troposphere as well as a cooling of up to 6 K in the NH high-latitudes. Enhanced cooling is also seen over Antarctica. The global-mean surface air temperature is 10.0 °C in LGM<sub>PMIP3</sub> and 10.5 °C in LGM<sub>PROXY</sub>, close to the "warm" LGM (10.7 °C) of Murray et al. [2014] which was based on the SST reconstruction from CLIMAP [CLIMAP, 1976].The result of LGM<sub>PROXY</sub> is quite similar to LGM<sub>PMIP3</sub> and thus not shown here.

### 3.2 Water vapor and hydroxyl radical (OH) changes

Water vapor transported from the tropical troposphere and oxidation of CH<sub>4</sub> in the stratosphere are the two primary sources of water vapor in the stratosphere [e.g., LeTexier et al., 1988]. Water vapor transported upward across the tropical tropopause is largely determined by tropical cold point tropopause (CPT) temperatures [e.g., Randel et al., 2004]. Figure 2 shows the annual- and zonal-mean CPT temperature and the water vapor concentration at the CPT versus latitude from the 3 simulations (left panels) and the changes relative to the PI (right panels). As compared to the PI, the ice-age conditions induce a cooling of ~1.5 K in the CPT in the tropics and NH high-latitudes, and 2-4 K cooling at the SH high-latitudes (Fig. 2b). It is also interesting to notice decreased cooling between about 35-65° latitudes in both hemispheres, which may be due to the equatorward shift of subtropical jets in cold climate [Fu et al., 2006; Fu and Lin, 2011]. Relative to PI, the water vapor concentration at the CPT in the LGM is decreased by ~1 ppmv in the tropics and up to 6 ppmv at high latitudes (Fig. 2d).

The relative changes of stratospheric water vapor, OH and NO<sub>x</sub> for LGM<sub>PMIP3</sub> versus PI are shown in Fig. 3, in which the climatology of PI is superimposed as contours in ppmv, pptv, and ppby, respectively. Stratospheric water vapor tends to increase with height because of increasing CH<sub>4</sub> oxidation [LeTexier et al., 1988]. There is a local minimum of water vapor concentration at the south pole owing to extreme cold temperatures there and associated dehydration through cloud particle sedimentation. Relative to PI, stratospheric water vapor in the LGM decreases everywhere, with a 10-25% decrease in the lower and middle stratosphere and a ~20% decrease in the upper stratosphere, as a consequence of much lower atmospheric CH<sub>4</sub> concentrations (Table 1), and the colder tropical CPT (Fig. 2b) leading to less water vapor transported from the troposphere (Fig. 2d). OH is mainly produced by the reaction of excited atomic oxygen, O(1D), with water vapor. In line with the decrease in stratospheric water vapor, the stratospheric OH decreases by 10-20% over the bulk of the stratosphere during the LGM. The detailed pattern of OH changes (Fig.3b) is determined by the changes of both H<sub>2</sub>O (Fig. 3a) and O(<sup>1</sup>D), and the latter is governed largely by the change in UV radiation. The reaction of N<sub>2</sub>O with O(<sup>1</sup>D) is the major source of stratospheric NO<sub>x</sub>. Because of much higher UV radiation at higher altitudes, the climatological NO<sub>x</sub> concentrations tend to increase with height, maximizing around 2 hPa (Fig. 3c). Due to the much lower N<sub>2</sub>O concentration in the LGM, the NO<sub>x</sub> mixing ratio decreases by 10-20% in the stratosphere, with small increases around the south pole and 100 hPa. The detailed pattern of NO<sub>x</sub> changes (Fig.3c) is again determined by the changes of both N<sub>2</sub>O and O(<sup>1</sup>D), and the small increases of NO<sub>x</sub> concentrations around the south pole and 100 hPa may be due to increases of  $O(^{1}D)$  in those regions (not shown).

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Compared with LGM<sub>PMIP3</sub>, more water vapor is transported from the troposphere into the stratosphere in LGM<sub>PROXY</sub> because of the warmer tropical SST. Both stratospheric water vapor and

OH in LGM<sub>PROXY</sub> are thus a bit higher than LGM<sub>PMIP3</sub>, but the changes relative to PI remain similar (not shown). For brevity, unless otherwise indicated, the results of LGM<sub>PROXY</sub> are not shown in the following discussions.

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### 3.3 Ozone change

Figure 4 shows the ozone change for LGM<sub>PMIP3</sub> versus PI. Relative to the PI, ice-age conditions lead to ozone decreases in most of the stratosphere, except for the tropical lower stratosphere where the ozone increases by 5-15%, extending up to about 20 hPa. The increase of ~5% near 10 hPa is caused by the maximum reductions in OH and NO<sub>x</sub> there (Fig. 3) while the increase around 100 hPa over the SH high-latitudes is linked to the local intensification of BDC in the spring season [see Fig. 3e in Fu et al., 2020]. The increases near ~300 hPa in both NH and SH extratropics are caused by lower tropopauses in the LGM which will be discussed latter. The increase in tropical lower stratospheric ozone is due to a decrease in the BDC in the LGM [Fu et al., 2020] and will be further discussed later. Compared to PI, major catalysts of stratospheric ozone destruction including both OH and NO<sub>x</sub> decrease in the LGM (Fig. 3), which will act to increase stratospheric ozone loss rates. On the other hand, owing to the reduced greenhouse gas concentrations, the stratosphere warms by up to 8 K in the LGM (Fig. 1). It is well known that stratospheric ozone in the upper stratosphere has an inverse relationship with temperature linked to gas-phase kinetic factors for several reactions displaying an energy of activation. For example, as shown in Jonsson et al. [2004], in response to doubled CO<sub>2</sub>-induced stratospheric cooling, stratospheric ozone increases because of slowing of the Chapman reaction  $O + O_3 \rightarrow 2O_2$  and thus slower ozone destruction rates. A warmer LGM stratosphere, on the other hand, leads to faster ozone loss reaction rates due to temperature-sensitive reactions which act to lower ozone

concentrations. The net result of these compensating effects is only a 5-10% decrease in upper stratospheric ozone despite large changes in stratospheric temperatures, generally consistent with previous LGM studies [Crutzen and Brühl, 1993; Martinerie et al., 1995].

234 Unlike the upper stratosphere, in which ozone photochemistry dominates, the BDC plays 235 a significant role in determining the lower stratospheric ozone distribution. Following Li et al. [2009], the mean advective ozone transport  $-(\overline{w^*} \frac{\partial \overline{O_3}}{\partial z} + \overline{v^*} \frac{\partial \overline{O_3}}{\partial y})$  in the model is presented in Fig. 236 5. Here,  $\overline{w^*}$  and  $\overline{v^*}$  are the transformed Eulerian mean (TEM) residual vertical and meridional 237 velocity [Andrews et al., 1987], and  $\overline{O_3}$  is the zonal mean ozone mixing ratio. The vertical 238  $(-\overline{w}^*\frac{\partial \overline{O_3}}{\partial z})$  and meridional  $(-\overline{v}^*\frac{\partial \overline{O_3}}{\partial y})$  components of advective ozone transport are also shown 239 240 (middle and far-right panels of Fig. 5). In the lower stratosphere (below ~15 hPa), the advective 241 ozone transport is negative in the tropics but positive in the extratropics. By contrast, there is 242 positive (negative) advective ozone transport in the tropics (extratropics) in the upper stratosphere. 243 Generally, the advective ozone transport pattern is dominated by the vertical component because 244 of much larger ozone vertical gradients compared to those in the horizontal (see contours of Fig.4). 245 It is interesting to note that meridional ozone transport tends to increase ozone everywhere, except 246 in the subtropical jet regions around the tropopause. The vertical and meridional advective ozone 247 transport are stronger in the NH than the SH, reflecting the stronger BDC there [Rosenlof, 1995].

Fu et al. [2020] found that the BDC in the LGM is slower than that in the modern climate, regardless of SSTs prescribed. Consistent with the weakening BDC in the LGM, the change of advective ozone transport (Figs. 5d-5f) is similar to, but generally the reverse sign of, the climatological pattern (Figs. 5a-5c). The positive change of ozone advective transport in the tropical lower stratosphere (Fig. 5d) explains the increase of ozone there in the LGM as already noted above (Fig. 4). The ozone advective lifetime is comparable with the chemical loss lifetime

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around 70 hPa but becomes much longer in the upper stratosphere [Li et al., 2009]. Hence, although LGM changes to advective ozone transport act to increase ozone in the upper stratosphere in the middle and high latitudes, chemical processes strongly dominate in this region, resulting in ozone decreases there.

In order to further quantify the changes of ozone in the LGM, the total ozone column is divided into three parts: the upper stratospheric column (above 15 hPa) where chemical processes dominate, the lower stratospheric column (below 15 hPa but above the tropopause) where dynamical transport dominates, and the tropospheric column (below the tropopause). Based on *in situ* trace gas measurements, Pan et al. [2018] illustrated that the World Meteorological Organization lapse rate tropopause (LRT) [World Meteorological Organization, 1957] did better than the CPT in identifying the transition from the tropopause to the stratosphere. Therefore, the partitioning between the stratospheric and tropospheric ozone column is based on LRT, namely the lower boundary of a layer in which the temperature lapse rate is less than 2 K km<sup>-1</sup> for a depth of at least 2 km.

Figure 6 shows the annual- and zonal-mean latitudinal distributions of partial ozone columns (upper panels) and their changes relative to the PI (lower panels). Since the lifetime of ozone related to chemical processes is much shorter than the timescale of advective transport in the upper stratosphere and the rate of ozone production is larger when the solar angle is smaller [Brasseur and Solomon, 2006], the upper stratospheric ozone column peaks in the tropics and decreases toward the high latitudes (Fig. 6a). In contrast, the lower stratospheric ozone column (Fig. 6b) maximizes in the high latitudes, with a minimum at the tropics, mainly because of advective ozone transport. As expected, most of the ozone resides in the stratosphere, and tropospheric ozone (Fig. 6d) is only a small proportion (~10%) of the total ozone column, with

higher concentrations in mid-latitudes and higher concentrations in the NH than the SH. This results from higher photochemical ozone production related to higher natural continental emissions of CO, NO<sub>x</sub> and hydrocarbons in the NH relative to the SH [e.g., Logan, 1985; Lelieveld and Dentener, 2000].

Due to the compensating effects of changes in major catalysts of ozone destruction (OH and NO<sub>x</sub>) and temperature, ice-age conditions lead to a small (~1-3 DU) decrease of the upper stratospheric ozone column relative to PI (Fig. 6f). Because of the weakening BDC, the lower stratospheric ozone column increases by about 2-5 DU in the tropics and decreases by 5-10 DU in the extratropics (Fig. 6g). The modeled change in the total stratospheric ozone column (Fig. 6h) is dominated by change in the lower stratospheric ozone column (Fig. 6g). A reduced decrease (i.e., an increase) of lower-stratospheric ozone (Fig. 6g) in the NH (SH) high latitudes is partly because of a lower tropopause there in the LGM as shown later.

As compared with PI, the tropospheric ozone column decreases by ~ 5DU (30%) in the LGM (Fig. 6i), which is within the range reported by previous modeling studies (see Table 1 summarized by Murray et al. [2014]). The ice-age conditions lead to decreases (5-15 DU) of total ozone column at almost all latitudes (Fig. 6j). The latitudinal dependence of the total ozone column change is dominated by the lower stratospheric ozone column change (Fig. 6g) while the tropospheric change shifts the total change down by about 5 DU. Near the equator and south pole the stratospheric and tropospheric ozone column changes tend to cancel each other.

Figure 7 shows the spatial distribution of stratospheric ozone column in the PI and change for LGM versus PI. As compared with PI, stratospheric ozone increases slightly in the tropics but decreases by up to 25 DU in mid- and high latitudes. The larger decrease of the stratospheric ozone column in the NH relative to the SH is due to the larger magnitude decrease of the BDC in the NH

[Fu et al., 2020]. It is interesting to note that a significant increase (~25 DU) of the stratospheric ozone column is found in northeast North America, which is collocated with the maximum LRT pressure increase (Fig. 8) and a strong local intensification of downward vertical motion (not shown). Those changes are probably driven by the presence of the Laurentide ice sheet.

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In order to quantify the relative contributions of changes in the height of the tropopause versus changes in transport on the lower stratospheric ozone column change in the LGM, we integrate the LGM stratospheric ozone field vertically but impose the PI tropopause. stratospheric ozone column change arising from the tropopause change is approximated by the difference between the result using the LGM ozone field combined with the PI tropopause and the result using the LGM ozone and tropopause fields. By subtracting the contribution of tropopause change from the lower stratospheric ozone column change, the residual is referred to as the transport contribution. Figure 9 shows the spatial distributions of upper and lower stratospheric ozone column changes for LGM<sub>PMIP3</sub> versus PI, and the contributions to the latter from the tropopause change and transport change. Consistent with Figure 6, the stratospheric ozone column change in the LGM (Fig. 7b) is dominated by the lower stratospheric ozone change (Fig. 9b). Further, we note that the contribution from tropopause changes leads to an increase of stratospheric ozone column in the extratropics (Fig. 9c), which can also be seen in Fig.4 near 300 hPa. Although these changes are generally much smaller than those from transport changes (Fig. 9d), they do partly compensate for the ozone decreases due to the circulation changes in the extratropics, especially in the SH (Figs. 9 and 6).

Figure 10 is the same as Figure 7, but for the tropospheric ozone column. Relative to the PI, the tropospheric ozone column in the LGM is decreased (~5 DU) everywhere, maximized around North America, which may be due to the decrease of biogenic emissions and much higher

tropospheric photolysis (Figs. 13b) caused by the ice sheets. Here, the tropospheric ozone change for the LGM versus PI is consistent with theoretical expectation based on temperature-dependent surface emission of ozone precursor gases, and does not support the hypothesis of Geng et al. [2017] that the transport of stratospheric ozone to the surface, driven by a stronger BDC, increases in the glacial climate.

Figure 11 is the same as Figure 7, but for the total ozone column. Compared to PI, the total ozone column in the LGM is decreased almost everywhere, except at the Northeast of North America, equatorial Indian and west Pacific Oceans, and Antarctic between 60 and 150° E, because of larger stratospheric ozone increases (Fig. 7b) relative to the tropospheric ozone decreases (Fig. 10b) in those regions.

## 3.4 UV radiation change

Ozone is one of the major factors determining surface UV radiation [Lamy et al., 2019], which in turn drives photochemical reactions in the troposphere. In this subsection, we investigate the change of photolytic rate constant ( $J_{O1D}$ ) of  $O_3$ + hv  $\rightarrow O_2$  +  $O(^1D)$ , which is driven by photons at wavelengths shorter than 320 nm. Figure 12 shows the annual- and zonal-mean  $J_{O1D}$  (day<sup>-1</sup>) change for LGM versus PI. Because of the absorption of UV radiation by ozone, climatological  $J_{O1D}$  tends to decrease toward the surface, with the largest values in the tropics due to low ozone columns and small solar angles. Compared to the PI,  $J_{O1D}$  in the LGM decreases in the tropics and SH high latitudes, extending from 50 hPa toward the surface, but increases in other regions. The decreases of  $J_{O1D}$  for the LGM versus PI in the tropics and SH high latitudes are due to the increases of stratospheric ozone concentrations there (Fig. 4), and the increases in other regions are caused by the decreases in total ozone column abundance (Fig. 11b). It is also noted that there are

enhanced increases in  $J_{O1D}$  in the NH and SH mid-latitudes near ~700 hPa.

Figure 13 shows the changes in the spatiotemporal distribution of surface J<sub>OID</sub> in the LGM. The seasonal cycle of surface J<sub>OID</sub> follows the seasonal progression of the solar zenith angle (Fig. 13a). A maximum of surface J<sub>O1D</sub> occurs over the Sahara Desert, which is related to the high surface albedo there. Higher J<sub>OID</sub> also occurs in regions with high topography, e.g. Tibet and Andes, because of higher elevations and thus smaller solar attenuation by the atmosphere. In line with the change of total ozone column (Fig. 11b), surface J<sub>OID</sub> in LGM decreases in the tropics and the Antarctic, but increases in most of other regions, with the largest increase in North America and Europe in the boreal summer season caused by the increase in surface albedo and smaller solar attenuation resulting from the large ice sheets there. A large increase of surface J<sub>O1D</sub> is also found over the SH near 60° S, which is due to increased sea ice extent in the LGM. The relative change of surface J<sub>O1D</sub> for LGM versus PI is shown in Fig. SM1. Interesting to note that, large relative changes (up to 100%) of the surface J<sub>O1D</sub> exist in high latitudes of NH in the LGM, because of small climatological values there (See Fig. 13a). The cloud fraction decreases in the high latitudes of the NH (Fig. SM2) may also make some contributions to the large increase of surface J<sub>OID</sub> there in the LGM, in addition to the total ozone column change (Fig. 11b). On the other hand, an increase in clouds in the midlatitudes (Fig. SM2), especially in the SH, leads to a decrease in J<sub>OID</sub> there, partly compensating the effect of total ozone column change.

Figure 13 indicates that the ice sheets have a large local impact on surface  $J_{O1D}$  change in the LGM. However, in regions not impacted by ice sheets (e.g., the tropics), the overhead ozone column abundance dominates the LGM change in  $J_{O1D}$ , resulting in decreases in photolysis rates in the tropics but increases in most other areas, as compared with the PI.

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### 4. Conclusion and Discussion

We have investigated stratospheric ozone in the LGM using WACCM6. It is shown that, compared to the PI, water vapor in the LGM decreases throughout the stratosphere, with a  $\sim$ 10-25% decrease in the lower and middle stratosphere, and a 20% decrease in the upper stratosphere. This is largely owing to the much lower atmospheric CH<sub>4</sub> concentration, colder CPT ( $\sim$ 1.5 K in the tropics) and less water vapor ( $\sim$ 1 ppmv) at the CPT. In line with the lower stratospheric water vapor and N<sub>2</sub>O concentrations in the LGM, stratospheric OH and NO<sub>x</sub> are both decreased by 10-20%, which acts to decrease stratospheric ozone loss rates. On the other hand, because of the lower greenhouse gas concentrations, the stratospheric temperature increases by up to 8 K in the LGM. This results in an increase in the ozone loss rate linked to temperature-sensitive gas-phase reactions, especially a faster Chapman reaction O + O<sub>3</sub>  $\rightarrow$  2O<sub>2</sub> [Jonsson et al., 2004]. The net result of these compensating effects is a 5-10% (1-3 DU) decrease in the upper stratospheric ozone concentration, generally consistent with previous studies [Crutzen and Brühl, 1993; Martinerie et al., 1995].

The BDC plays a significant role in determining the spatial distribution of ozone in the lower stratosphere. Contrary to Rind et al. [2001] and Rind et al. [2009], who found a stronger BDC in the lowermost stratosphere in the LGM, Fu et al., [2020] showed that the BDC in the LGM is slower than the modern climate throughout the stratosphere. Because of the weakening of the BDC in the LGM, lower stratospheric ozone increases by 2-5 DU in the tropics and decreases by 5-10 DU in the extratropics. We also examined the effect of lower tropopause in the LGM, which increases the lower-stratospheric ozone in the extratropics and partly compensates the ozone decrease there due to the BDC changes.

In our simulations, we found that tropospheric ozone decreases by about 5 DU (30%) in the LGM, as compared with the PI. Such a result is consistent with theoretical expectation based

on temperature-dependent surface emissions of ozone precursor gases. As shown in Murray et al. [2014], transport of ozone from the stratosphere becomes an increasingly important fraction of the total tropospheric ozone in the cold climate, compared to the modern climate. Hence, changes in stratosphere-troposphere ozone exchange may play an important role in tropospheric chemistry in the LGM. The relative contributions of ozone formed *in situ* in the troposphere versus transport from the stratosphere in the LGM, and its implications for tropospheric chemistry, will be examined in a follow up paper.

It is interesting to note that stratospheric ozone increases significantly in the northeast of North America in the LGM, which is attributed to the local intensification of BDC induced by the Laurentide ice sheet. Compared to PI, the total ozone column in the LGM decreases everywhere, except over the northeast of North America, equatorial Indian and west Pacific Oceans, and East Antarctica, because of the larger stratospheric ozone increases relative to tropospheric ozone decreases there.

Compared to PI, surface J<sub>O1D</sub> in the LGM increases almost everywhere due to the decreased ozone column abundance. The largest increases are over North America and Europe because of the increase of surface albedo and smaller solar attenuation resulting from the large ice sheets. The large increase of surface J<sub>O1D</sub> over the SH near 60° S is caused by the increased sea ice extent in the LGM. Surface J<sub>O1D</sub> decreases in the tropics and the Antarctic in the LGM are due to the increases in the overhead ozone column abundance there. Such UV radiation change may have important ramifications for tropospheric chemistry [e.g. Murray et al., 2014], with potential implications for the methane lifetime as well as for interpretation of some paleoclimate records such as pollen [Benca et al., 2018]. A dynamically consistent ozone field in the LGM can also be expected to be important for model simulations of the LGM climate [e.g. Nowack et al. 2015;

Chiodo and Polvani, 2017; Noda et al., 2017; Szopa et al. 2019]. Therefore, in addition to the implications for tropospheric chemistry, the ozone fields from this study are useful for the PMIP LGM climate simulations. Acknowledgements: This research was supported by NSF Grant AGS-1821437 and NASA Grant 80NSSC18K1031. SS was supported by NSF AGS project 1848863. RHW was partially funded by the European Union's Horizon 2020 research and innovation programme under the Marie Sklodowska-Curie Grant Agreement No. 797961. We would like to acknowledge high-performance computing support from Cheyenne (doi:10.5065/D6RX99HX) provided by NCAR's Computational and Information Systems Laboratory, sponsored by the National Science Foundation, for the WACCM6 simulations and analyses presented in this study and for the data management, storage and preservation. We intend to provide access to the data by acceptance. The National Center for Atmospheric Research is sponsored by the United States National Science Foundation. 

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- 576 Figure captions577 Figure 1. Annu
- Figure 1. Annual-mean zonal-mean temperature change (color scale) for LGM<sub>PMIP3</sub> versus PI.
- Black contours show the annual-mean climatology (K) of PI.
- Figure 2. Annual-mean zonal-mean a) cold point tropopause (CPT) temperature and c) water vapor
- 580 (H<sub>2</sub>O) concentration at the CPT in 3 simulations. Panels b) and d) are the changes of CPT
- temperature and H<sub>2</sub>O concentration at the CPT relative to PI.
- Figure 3. Annual-mean percentage changes of a) water vapor (H<sub>2</sub>O), b) hydroxyl radical (OH),
- and c) nitrogen oxides (NO<sub>x</sub>) for LGM<sub>PMIP3</sub> versus PI. Black contours show the climatology of
- PI expressed in a) ppmv, b) pptv and c) ppbv, respectively.
- Figure 4. Annual-mean percentage change of ozone (color scale) for LGM<sub>PMIP3</sub> versus PI. Black
- contours show the climatology (ppmv) of PI.
- Figure 5. Advective ozone transport (far-left), and its vertical (middle) and meridional (far-right)
- components in the climatology of PI (upper panels) as well as the changes in LGM<sub>PMIP3</sub> versus
- 589 PI (lower panels).
- Figure 6. Annual-mean zonal-mean column ozone concentrations for various atmospheric layers
- (upper panels) and the changes relative to PI (lower panels).
- 592 Figure 7. Spatial distribution of stratospheric ozone column in a) PI simulation, and b) the absolute
- change for LGM<sub>PMIP3</sub> versus PI.
- Figure 8. Spatial distribution of lapse-rate tropopause (LRT) pressure change for LGM<sub>PMIP3</sub> versus
- 595 PI.
- Figure 9. Spatial distribution of ozone column changes in the a) upper and b) lower stratosphere
- between LGM<sub>PMIP3</sub> and PI. Panels c) and d) are the lower stratospheric ozone column changes
- resulting from the tropopause and transport changes.

Figure 10. Same as Figure 7, but for tropospheric ozone column. Figure 11. Same as Figure 7, but for total ozone column. Figure 12. Annual-mean absolute change in photolysis rate constant (J<sub>O1D</sub>) for the reaction O<sub>3</sub> +  $hv \rightarrow O_2 + O(^1D)$  in LGM<sub>PMIP3</sub> versus PI. Black contours show the climatology (day<sup>-1</sup>) of PI. Figure 13. Spatiotemporal distribution of surface photolysis rate J<sub>O1D</sub> in a) PI simulation and b) the absolute change for LGM<sub>PMIP3</sub> versus PI. Each row contains two plots sharing a common ordinate axis (latitude). From left to right, a Hovmöller diagram presents the seasonality of zonal mean surface photolysis rate, and a map presents the horizontal distribution. 

# Table 1. Model Simulations.

Simulation names		$LGM_{PMIP3}$	$LGM_{PROXY}$	PI
$CO_2$		185 ppm	185 ppm	284 ppm
$\mathrm{CH_4}$		349 ppb	349 ppb	810 ppb
$N_2O$		200 ppb	200 ppb	273 ppb
ODS		0	0	0
СО	Fires	0.1 × Preind.	$0.1 \times Preind.$	Year 1850
	Ocean	Preind.	Preind.	Year 1850
	Fuel Combustion	0	0	Year 1850
	Lighting	Interactive	Interactive	Interactive
$NO_x$	Fires	$0.1 \times Preind.$	$0.1 \times Preind.$	Year 1850
NO <sub>x</sub>	Soil	$0.98 \times Preind.$	$0.98 \times Preind.$	Year 1850
	Fuel Combustion	0	0	Year 1850
SSTs		PMIP3* multi-model mean	MARGO proxy data	1870-1890 average
Sea ice		PMIP3 multi-model mean	PMIP3 multi-model mean	1870-1890 average
Ice sheet topography		Abe-Ouchi et al. [2015]	Abe-Ouchi et al. [2015]	Modern climate
Orbital forcing		21,000 years ago	21,000 years ago	Year 1850

<sup>\*</sup> PMIP3 is Paleoclimate Modeling Intercomparison Project Phase 3.

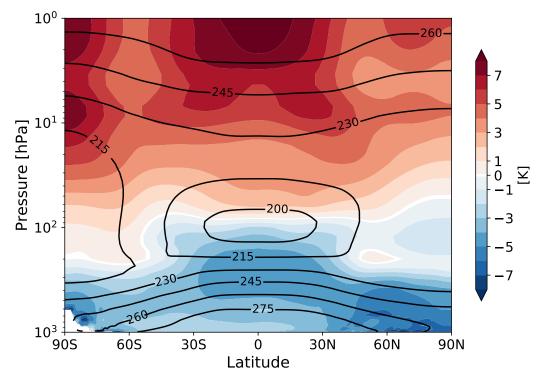


Figure 1. Annual-mean zonal-mean temperature change (color scale) for  $LGM_{PMIP3}$  versus PI. Black contours show the annual-mean climatology (K) of PI.

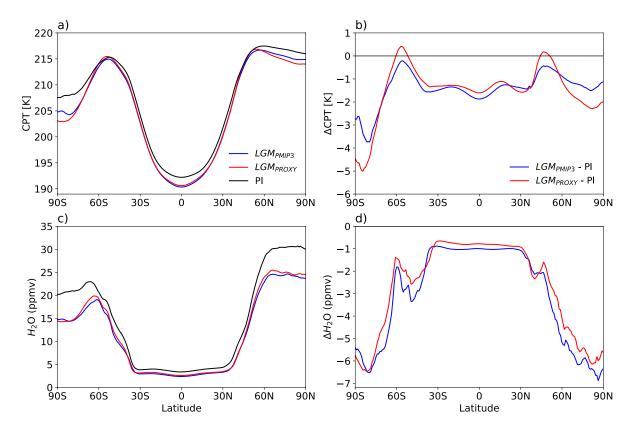


Figure 2. Annual-mean zonal-mean a) cold point tropopause (CPT) temperature and c) water vapor  $(H_2O)$  concentration at the CPT in 3 simulations. Panels b) and d) are the changes of CPT temperature and  $H_2O$  concentration at the CPT relative to PI.

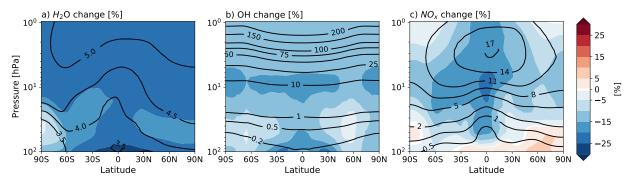


Figure 3. Annual-mean percentage changes of a) water vapor  $(H_2O)$ , b) hydroxyl radical (OH), and c) nitrogen oxides  $(NO_x)$  for  $LGM_{PMIP3}$  versus PI. Black contours show the climatology of PI expressed in a) ppmv, b) pptv and c) ppbv, respectively.

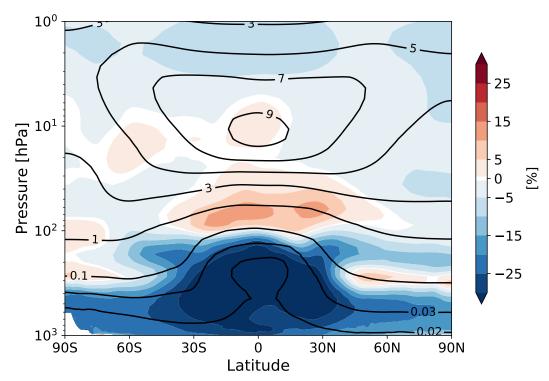


Figure 4. Annual-mean percentage change of ozone (color scale) for LGM<sub>PMIP3</sub> versus PI. Black contours show the climatology (ppmv) of PI.

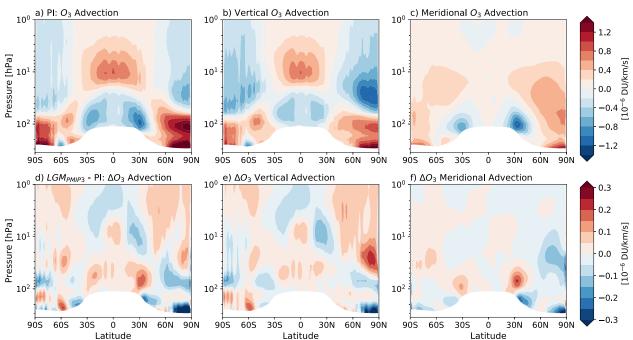


Figure 5. Advective ozone transport (far-left), and its vertical (middle) and meridional (far-right) components in the climatology of PI (upper panels) as well as the changes in LGM<sub>PMIP3</sub> versus PI (lower panels).

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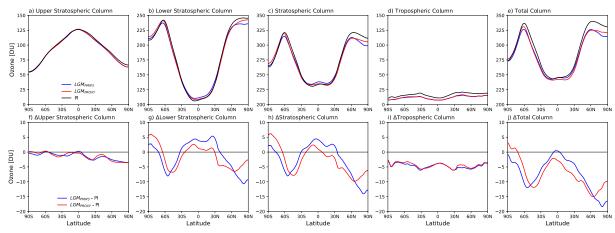


Figure 6. Annual-mean zonal-mean column ozone concentrations for various atmospheric layers (upper panels) and the changes relative to PI (lower panels).

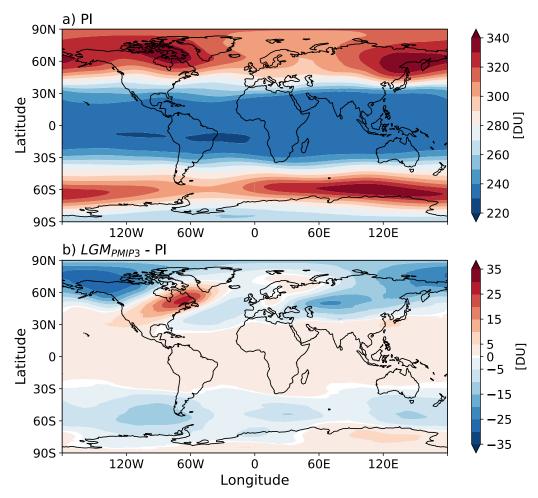


Figure 7. Spatial distribution of stratospheric ozone column in a) PI simulation, and b) the absolute change for LGM<sub>PMIP3</sub> versus PI.

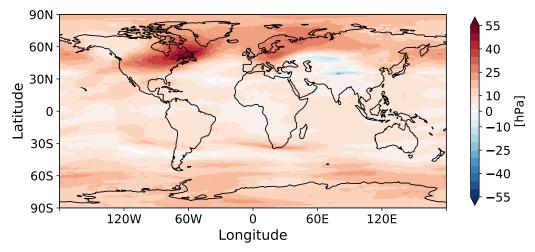


Figure 8. Spatial distribution of lapse-rate tropopause (LRT) pressure change for  $LGM_{PMIP3}$  versus PL

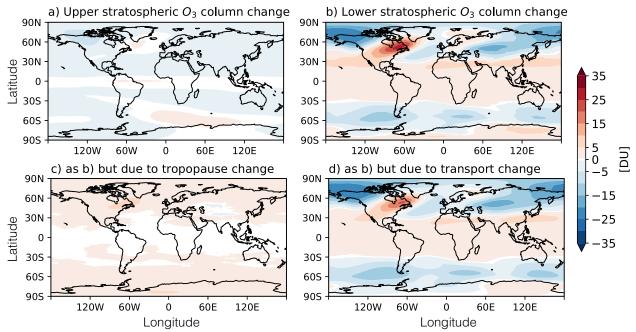


Figure 9. Spatial distribution of ozone column changes in the a) upper and b) lower stratosphere between LGM<sub>PMIP3</sub> and PI. Panels c) and d) are the lower stratospheric ozone column changes resulting from the tropopause and transport changes.

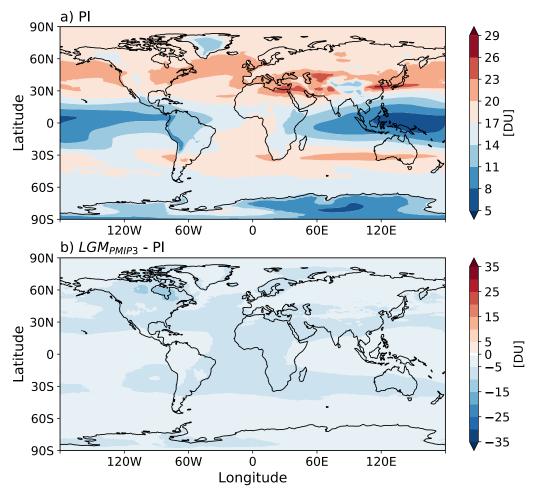


Figure 10. Same as Figure 7, but for tropospheric ozone column.

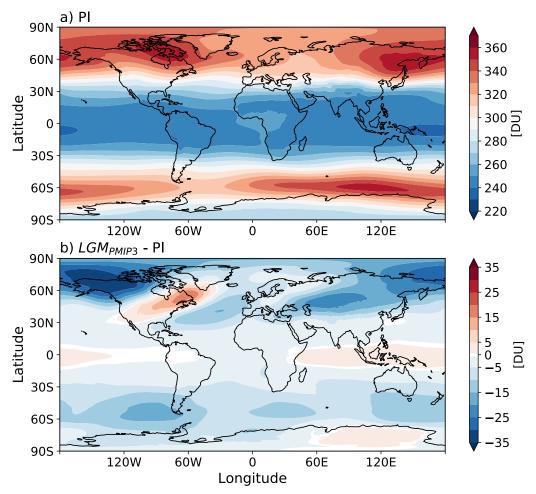


Figure 11. Same as Figure 7, but for total ozone column.

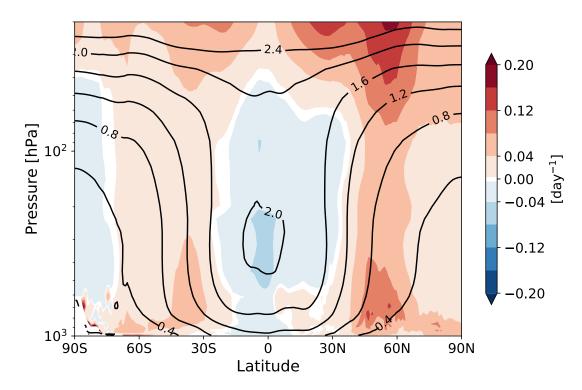


Figure 12. Annual-mean absolute change in photolysis rate constant  $(J_{O1D})$  for the reaction  $O_3$  +  $hv \rightarrow O_2 + O(^1D)$  in  $LGM_{PMIP3}$  versus PI. Black contours show the climatology (day<sup>-1</sup>) of PI.

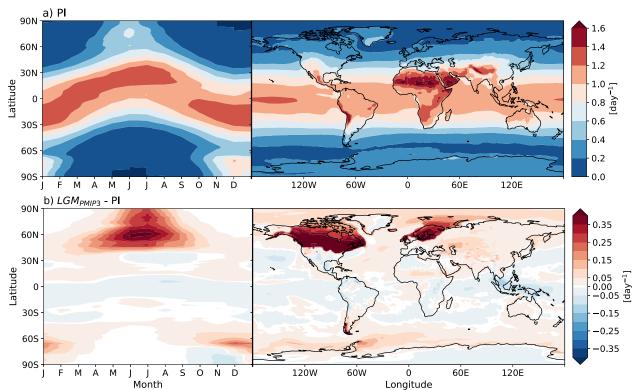


Figure 13. Spatiotemporal distribution of surface photolysis rate  $J_{O1D}$  in a) PI simulation, and b) the absolute change for LGM<sub>PMIP3</sub> versus PI Each row contains two plots sharing a common ordinate axis (latitude). From left to right, a Hovmöller diagram presents the seasonality of zonal mean surface photolysis rate, and a map presents the horizontal distribution.